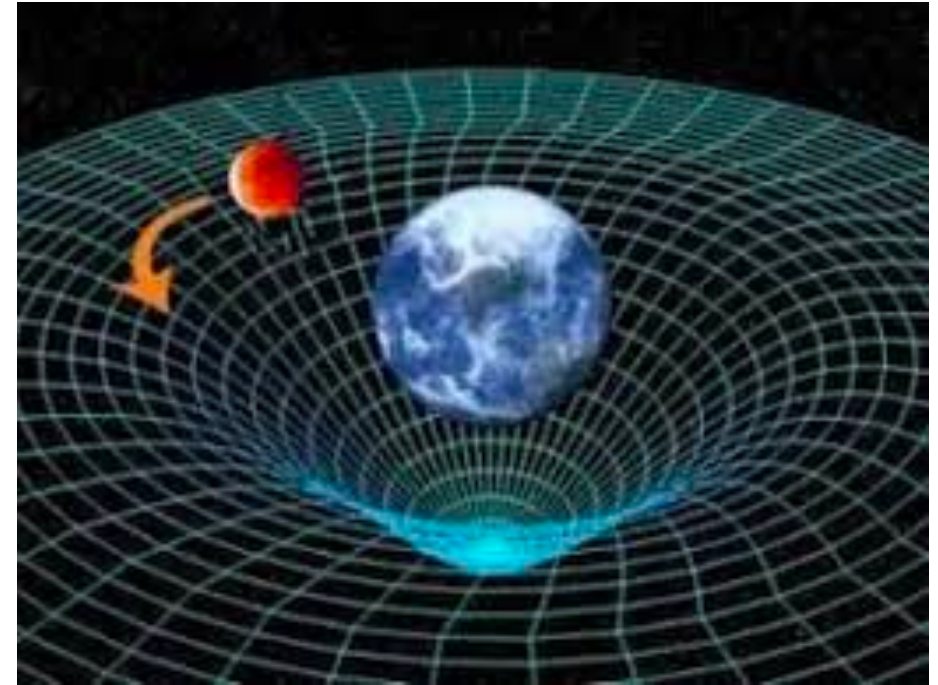


# GRAVITY AND SPACE



# Gravity

- Distortions of spacetime due to mass
- Mass doesn't weigh down spacetime, spacetime curves around a mass
- Attractive force



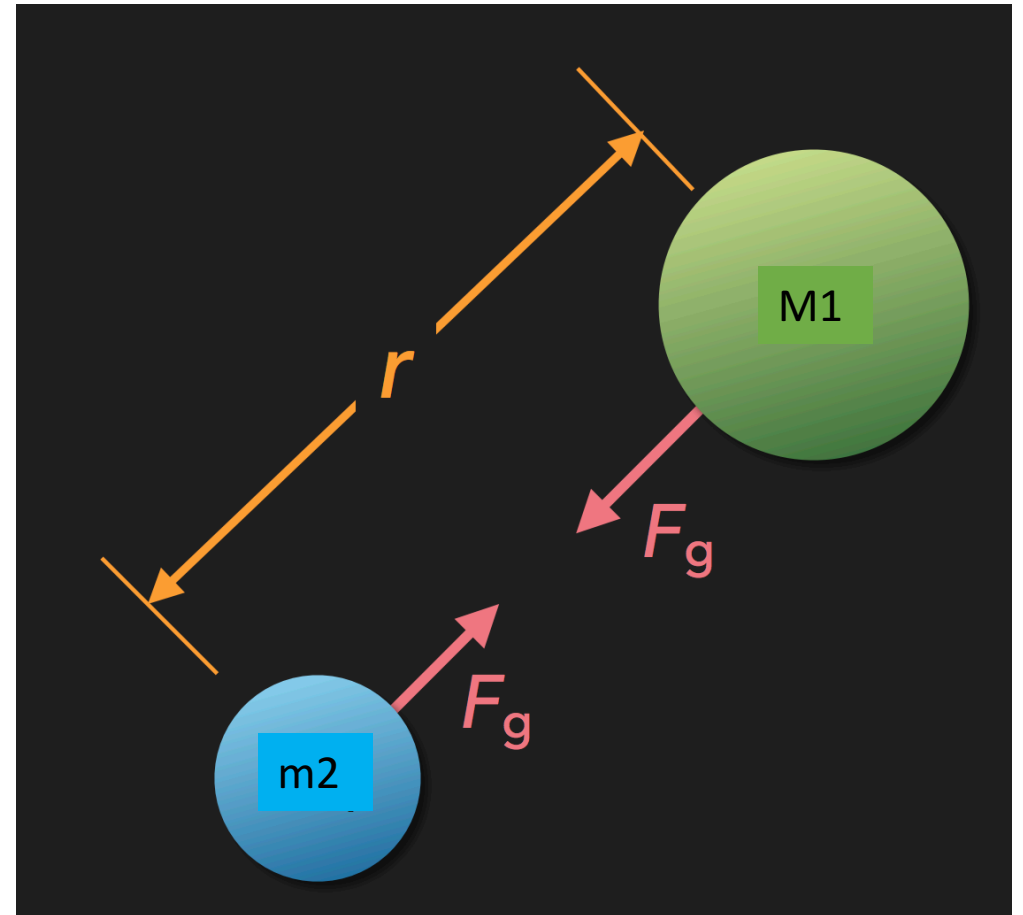
# Force of Gravity between two objects

$$F_G = \frac{GM_1m_2}{r^2}$$

$$G = 6.674 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$$

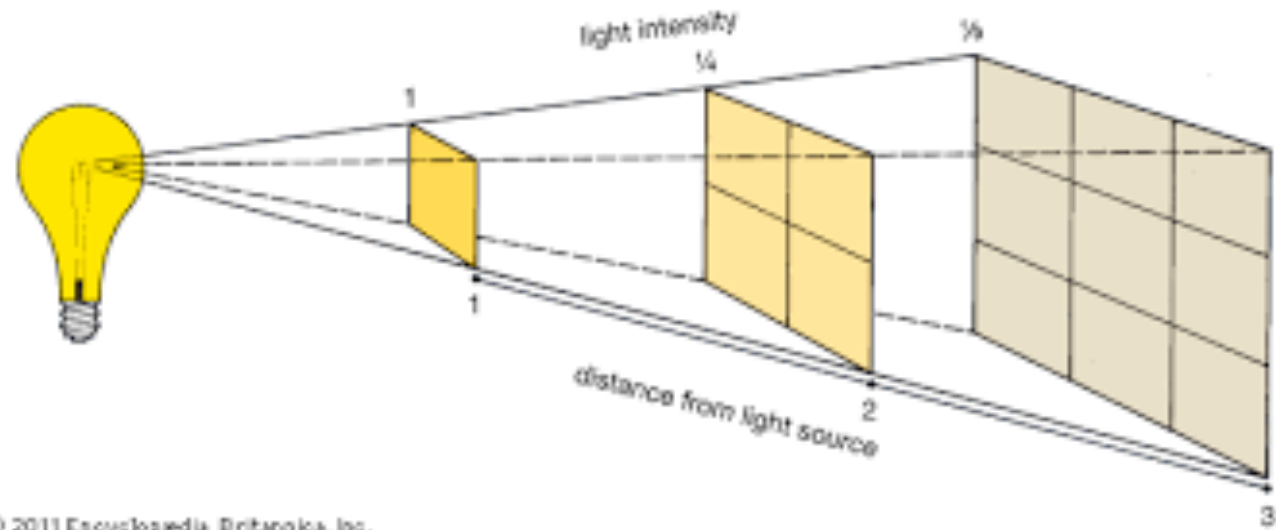
Universal gravitation constant  
 $r$  is distance between the **centers** of the objects

Which exerts a greater force:  
 $m_1$  on  $m_2$ , or  $m_2$  on  $m_1$ ? (Hint: think of Newton's 3<sup>rd</sup> Law)

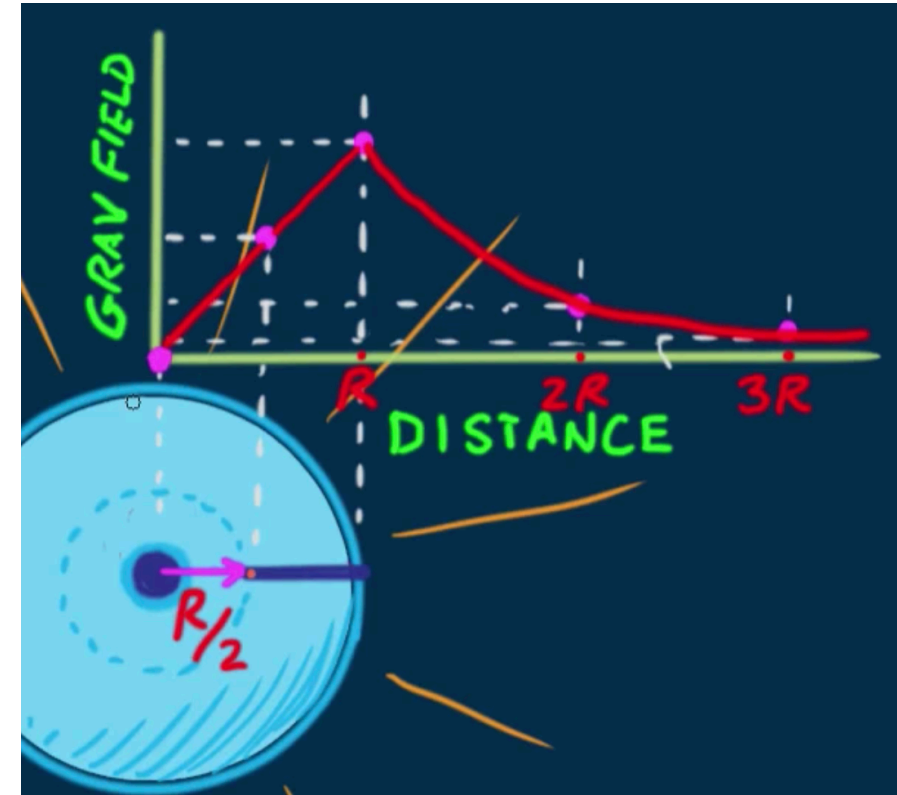


# Inverse square law

- The force of gravity is proportional to the inverse square of the distance
  - If you double the distance, the force of gravity decreases by a factor of 4.



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An apple on a tree feels 1 N of force due to gravity. If you double the height of the tree, the force of gravity on the apple would be:

- A. 2x as strong
- B.  $1/2$  as strong
- C.  $1/4$  as strong
- D. None of the above

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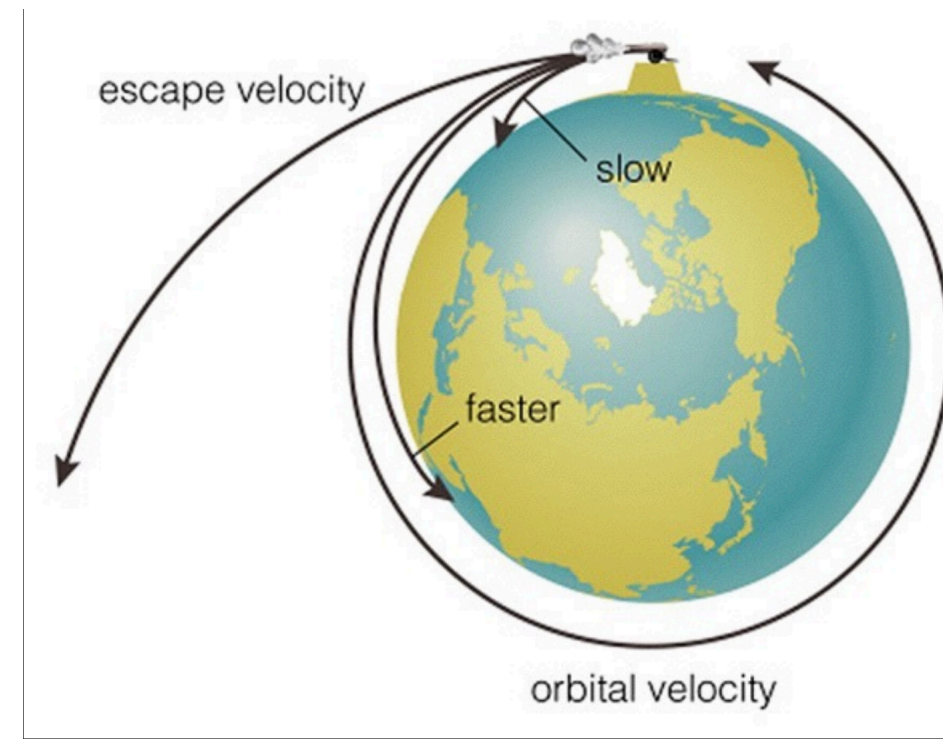
Why?

# Gravitational Potential Energy → Escape speed

- $PE_G = \frac{-GM_1m_2}{r}$
- How fast would you need to throw a ball of mass  $m$  from the surface of Earth for it to escape Earth's gravity?
- $E_{\text{tot}} = \text{KE} + \text{PE}$       When  $m$  is no longer in Earth's gravity well,  $E_{\text{tot}} = 0$
- $E_{\text{tot}} = \frac{1}{2}mv^2 - \frac{GM_E m}{r} = 0$
- $v_{\text{esc}} = \sqrt{\frac{2GM_E}{R_E}}$       doesn't depend on the mass of the ball!

# What is the escape speed of any object from the surface of Earth?

- $v_{esc} = \sqrt{\frac{2GM}{R}}$
- Mass of the Earth =  $5.98 \times 10^{24}$  kg, radius of Earth =  $6.38 \times 10^6$  m
- 11,200 m/s
  - Or 11.2 km/s



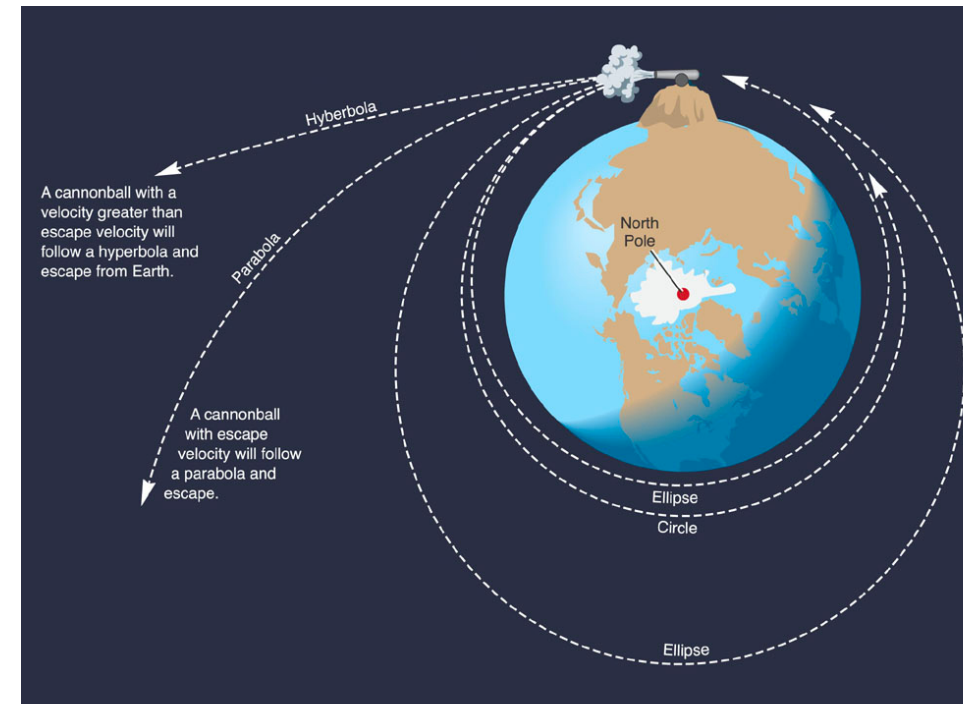


Which requires more fuel: a rocket going from Earth going to the Moon, or going from Moon to the Earth?

- Think escape velocity/which object has a larger gravity well

# Orbiting speed

- If you threw an object horizontally at 8000 m/s, it would travel 8000 m horizontally and fall 5m. It would never hit the ground and orbit the earth in a circular orbit
- What if you threw it at 9000 m/s? This is not quite fast enough to escape Earth's gravity well (escape speed = 11200 m/s)
  - It would orbit in an ellipse!



# “Weightlessness”

- Astronauts on ISS aren't really weightless
  - It's also a common misconception that there is no gravity in space
- They experience about 90% of the force of gravity from the Earth as we do
- They are constantly falling around the Earth



# Orbiting speed



- $F_G = F_c$

- $\frac{GM_1m_2}{r^2} = \frac{mv^2}{r}$

the  $m$  in the centripetal force equation is the one that is orbiting the earth, which gets cancelled out

- $v_{orb} = \sqrt{\frac{GM_1}{r}}$

# Orbiting speed

- If an object is orbiting at height  $h$  from the surface of the Earth, what is its orbiting speed?

- $v_{orb} = \sqrt{\frac{GM_1}{r}}$

- $r = R + h$

$R$  = radius of the Earth/object being orbited around



How does conservation of angular momentum explain why an object moves faster when it is closer to the sun vs. farther away from the sun?

- $L = I\omega$

If an asteroid was very small but super massive, could you really live on it like the Little Prince?



If the asteroid was 1.75 m in radius, what mass would it need to have in order to have Earthlike gravity at its surface?

- $g = 9.8 \text{ m/s}^2$

- $g = \frac{GM_1}{r^2}$

- $M = 4.5 \times 10^{11} \text{ kg}$

- That's just under 500 million tons, or the combined mass of every human on Earth...





# What would be the escape speed on the surface?

- $r = 1.75 \text{ m}$ ,  $M = 4.5 \times 10^{11} \text{ kg}$ ,

- $G = 6.674 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$

- $v_{esc} = \sqrt{\frac{2GM}{R}} = 5.9 \text{ m/s}$

- If you can't dunk a basketball, you can't escape by jumping, but you might escape by running and jumping off a ramp

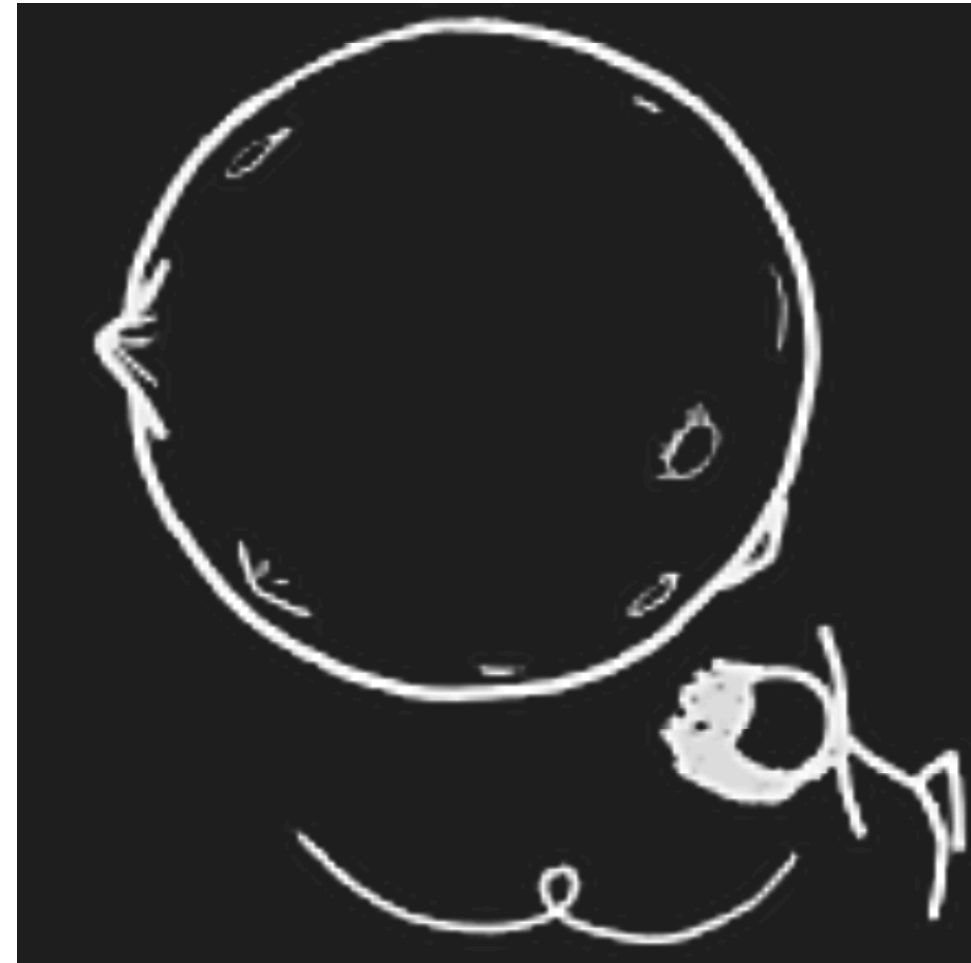


If your center of mass is 1.4 m above your feet, how fast would you need to run in order to orbit the asteroid?

- $r = 1.75 \text{ m}$ ,  $M = 4.5 \times 10^{11} \text{ kg}$ ,  $G = 6.674 \times 10^{-11} \text{ Nm}^2/\text{kg}^2$

- $v_{orb} = \sqrt{\frac{GM}{R+h}} = 3.1 \text{ m/s}$

- That's a typical jogging speed



# What would be the difference in the force of gravity at your head vs. your feet?

- Say you are 60 kg and 1.6 m (~5'3") tall
- $F_G = \frac{GM_1m_2}{r^2}$
- $F_{G \text{ at head}} = 160 \text{ N}$
- $F_{G \text{ at feet}} = 588 \text{ N}$
- Called **tidal forces** – when there's a substantial difference in the force of gravity at different ends of a body.
- What happens with Earth's tides and what we think will happen to you if you got too close to a black hole
- It would feel like lying on a merry-go-round with your head near the center



# Kepler's 3 Laws

- Johannes Kepler (1571-1630)
- German mathematician, astrologer, astronomer
- 3 Laws of Planetary Motion
  - 1) The orbit of a planet is an ellipse with the Sun at one of the two focus points 2) A line segment joining a planet and the Sun sweeps out equal areas during equal intervals of time 3) The square of the orbital period of a planet is proportional to the cube of its average distance from the Sun

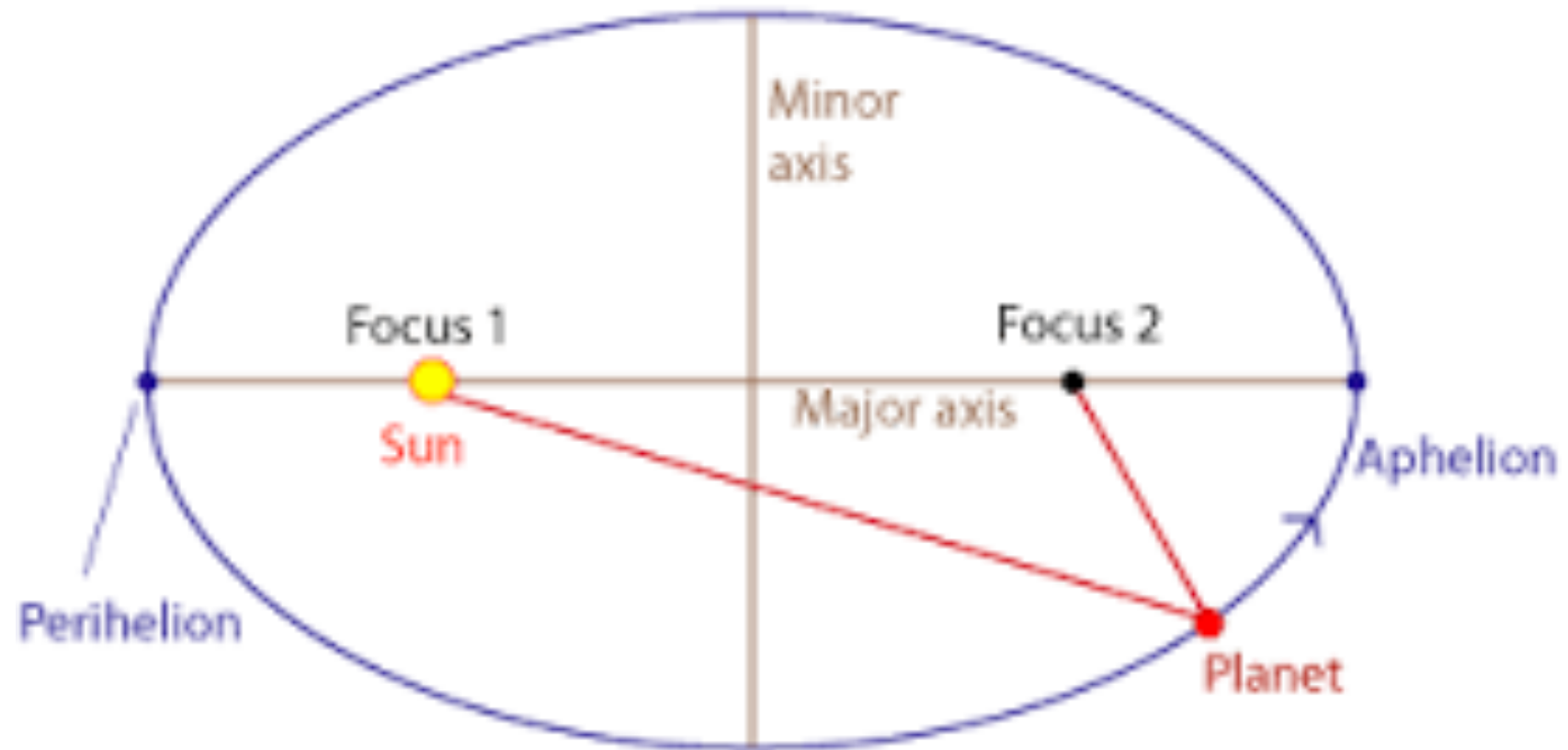


# Kepler's 3 Laws

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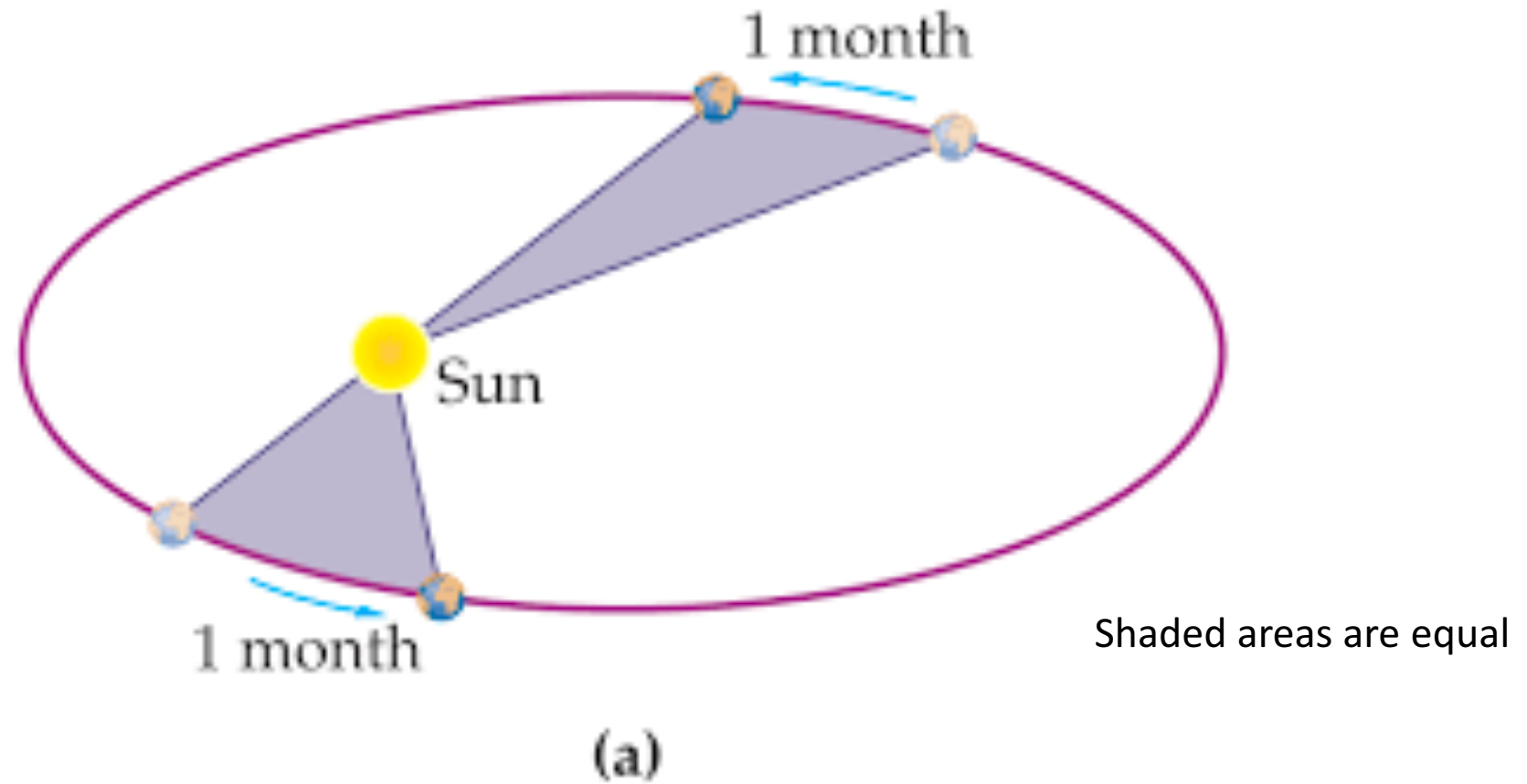


# Kepler's 1<sup>st</sup> Law



*An elliptical orbit of a planet  
(greatly exaggerated)*

# Kepler's 2<sup>nd</sup> Law



# Kepler's 3<sup>rd</sup> Law

$$v_{orb} = v_t$$

$$\sqrt{\frac{GM_1}{r}} = \frac{2\pi r}{T}$$

$$T^2 = \frac{4\pi^2}{GM} r^3$$

$$T^2 = r^3$$

If expressed in the following units:

$T$  Earth years

$r$  Astronomical units AU  
( $a = 1$  AU for Earth)

$T$  is in seconds and  $r$  is in meters

$r$  is the distance between the centers of the two objects, not the radius of the object!



Use Kepler's 3<sup>rd</sup> Law to find the mass of the Sun, if Earth's period is 1 year =  $3.154 \times 10^7$  seconds and the distance between the Earth and the sun is  $1.5 \times 10^{11}$  m

$$T^2 = \frac{4\pi^2}{GM} r^3$$

Solve for M:  
 $2.0 \times 10^{30}$  kg

